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Relaxation of N-body systems with additive $r^{-\alpha}$ interparticle forces

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ABSTRACT

In Newtonian gravity the final states of cold dissipationless collapses are characterized by several structural and dynamical properties remarkably similar to those of observed elliptical galaxies. Are these properties a peculiarity of the Newtonian force or a more general feature of long-range forces? We study this problem by means of N -body simulations of dissipationless collapse of systems of particles interacting via additive $r^{-\alpha}$ forces. We find that most of the results holding in Newtonian gravity are also valid for $\alpha \neq 2$. In particular the end products are triaxial and never flatter than an E7 system, their surface density profiles are well described by the Sérsic law, the global density slope-anisotropy inequality is obeyed, the differential energy distribution is an exponential over a large range of energies (for $\alpha \geq 1$), and the pseudo phase-space density is a power law of radius. In addition, we show that the process of virialization takes longer (in units of the system's dynamical time) for decreasing values of α , and becomes infinite for $\alpha = -1$ (the harmonic oscillator). This is in agreement with the results of deep MOND collapses (qualitatively corresponding to $\alpha = 1$) and it is due to the fact the force becomes more and more similar to the $\alpha = -1$ case, where as well known no relaxation can happen and the system oscillates forever.

Key words: gravitation – stellar dynamics – galaxies: kinematics and dynamics – methods: numerical

1 INTRODUCTION

One of the most striking properties of elliptical galaxies is the remarkable quasi-homology of their surface brightness profiles, described by the so called Sérsic model, a generalization of the de Vaucouleurs $R^{1/4}$ model (see e.g. Caon et al. 1993, Andrejakis et al. 1995, Courteau et al. 1996, Graham & Colless 1997, Prugniel & Simion 1997, Graham 1998, Trujillo et al. 2001, Bertin et al. 2002; see also Ciotti 2009 and references therein). Albeit minor (but important) departures from the Sérsic model are common, overall the profiles on large scale are very well represented by the Sérsic law. What is the origin of such regularity? N -body numerical simulations revealed that cold dissipationless and collisionless collapses lead to virialized end-states described almost perfectly by $R^{1/4}$ profiles (e.g. van Albada 1982, Londrillo et al. 1991). More recently it has been shown that collapses in pre-existing dark matter halos are also well described by the Sérsic profile, with a wide range of values of the Sérsic index (Nipoti et al. 2009a,b, hereafter N09a,b). In addition, it is also known that the Sérsic

family is characterized by an exponential differential energy distribution, over a large range of accessible energies (e.g. Binney 1982; Ciotti 1991). These results can be understood in terms of the physics of violent relaxation in collisionless collapses (e.g. Lynden-Bell 1967, Bertin & Stiavelli 1984, Bertin & Trenti 2003, Trenti & Bertin 2005, Trenti et al. 2005). Finally, it has been proved analytically that in Newtonian gravity a large class of spherically symmetric equilibrium systems are characterized by the so-called Global Density Slope-Anisotropy Inequality (hereafter GDSAI; see Ciotti & Morganti 2010a,b; van Hese et al. 2011, An et al. 2012; see also An & Evans 2006), a constraint between their anisotropy and density profiles. Numerical simulations suggest that the GDSAI may be a much more general result, holding true also for the final states of dissipationless collapses (see e.g. Hansen & Moore 2006).

Due to the relevance of these results for the understanding of the process of collisionless relaxation, a natural question arises about their apparent universality. In particular, are the Sérsic law, the associated differential exponential energy distribution and the GDSAI peculiar features of Newtonian gravity or are they more general properties of the

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